

Due to a special support structure the dishes surface keeps its parabolic form for all elevation positions but changes its focal lenght

Bild: Norbert Junkes (03.06.17)

The support structure





The dish has a surface of 9058 m<sup>2</sup>.

The outer area is transparent for short wavelengthes

# Receivers work down to a wavelength of 3mm

Bilder: esys.org; njn (12.04.14)





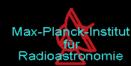


The basement must carry 3200 tonns

The valley shields man made radio radiation Valley open to the south To view the galactic center

**Bild: Norbert Tacken/MPIfR** 





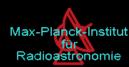


The telesecope runs on a Track, diameter 64 m. Centered by a vertical cylinder roller bearing

16 motors for azimuth, 2 in Elevation, each motor 17 KV

**Bilder: MPIfR** 



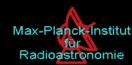


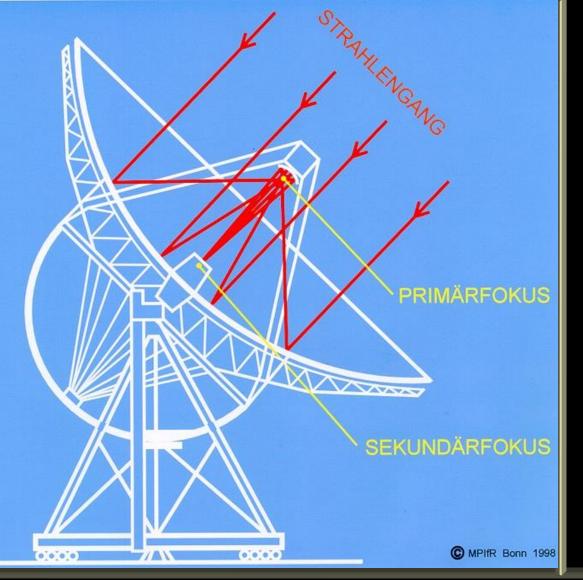


The cable twist with position encoders

**Bild: Norbert Tacken/MPIfR** 







Primary focus 30 m above dish surface

# Secondary mirror's diameter 6,5 m

Bild: MPIfR 1998





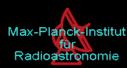


The primary focus receiver remote controlled motor driven forth and back to focus.

Retrack and close flaps for secondary foccus setup

Bilder: N. Tacken, K. Grypstra (MPIfR)

**Primary focus setup** 

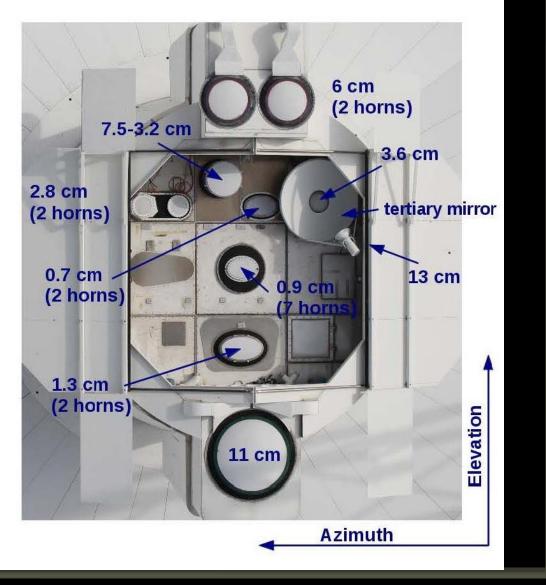




Bilder: njn (16.06.07)

Secondary focus setup





#### **Apex receivers**

Receivers range from 7 mm (Q-band: 33-50 GHz, since 03/2018) to 11 cm.

The horn size depends on the wavelength – the 11-cm-receiver horn's diameter is aprox. 1 m.

Bild: MPIfR (09/2018)

Secondary focus

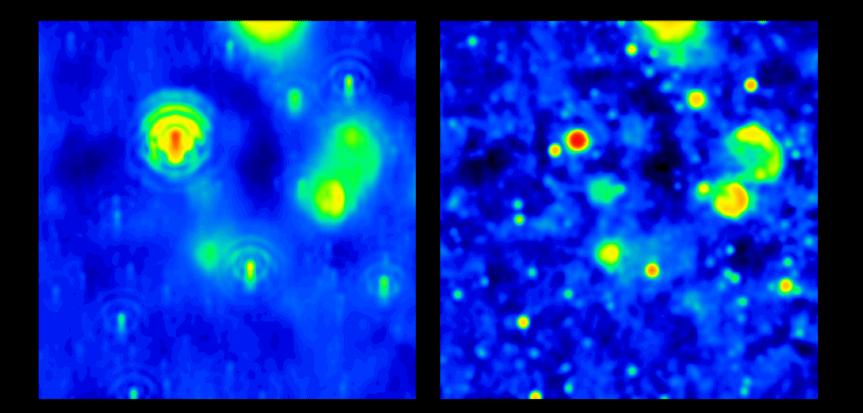






### Bilder: njn (05.10.06)

## The effect of active optics



### before

### afterwards

